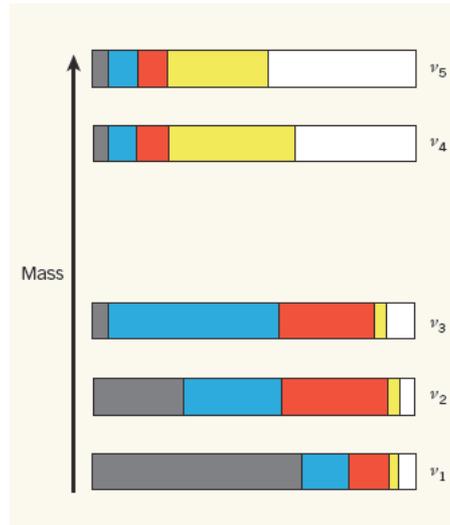


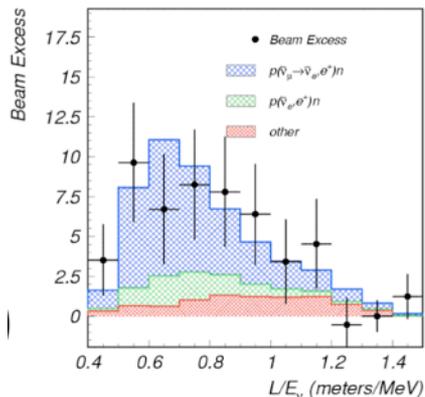
# Source and Reactor Experiments to Search for Sterile Neutrinos



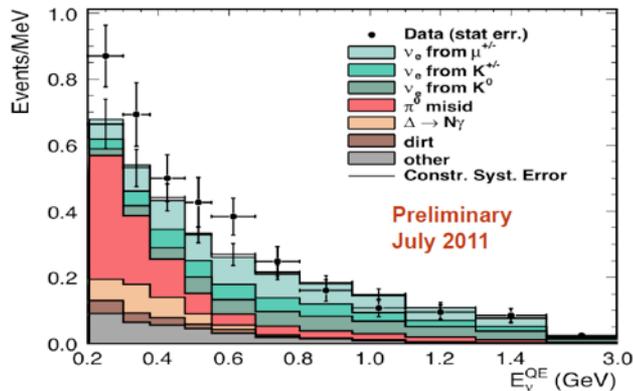
Karsten Heeger  
University of Wisconsin

# Neutrino Anomalies & Sterile $\nu$ Hypothesis

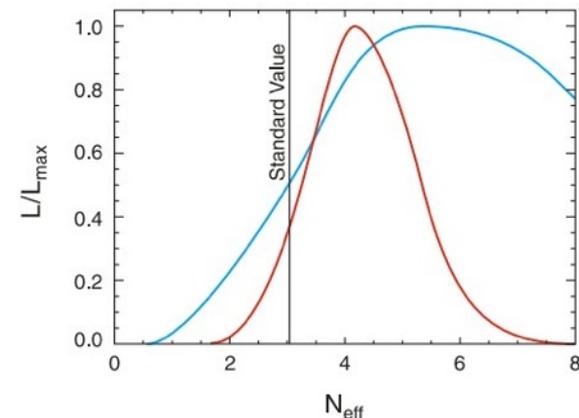
LSND



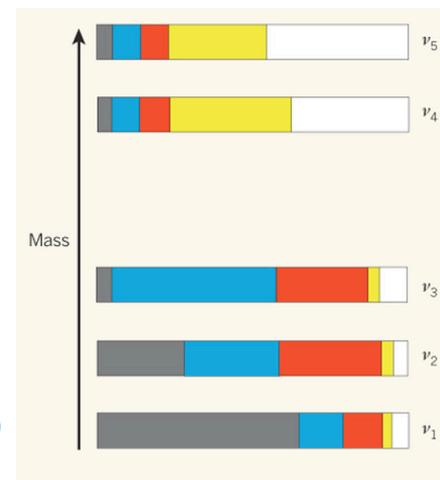
MiniBoone



Cosmology WMAP 7-year



- Experiments **at odds** with the standard **3-neutrino** interpretation of global neutrino oscillation data:
  - LSND ( $\bar{\nu}_e$  appearance)
  - MiniBoone anti-neutrinos ( $\bar{\nu}_e$  appearance)
  - Short baseline reactor experiments (re-evaluation of neutrino fluxes) ( $\bar{\nu}_e$  disappearance)
- If interpreted as oscillation signals  $\rightarrow$  a 4th (or more) **sterile neutrino** with  $\Delta m^2 \sim O(1 \text{ eV}^2)$  and  $\sin^2 2\theta > 10^{-3}$ .

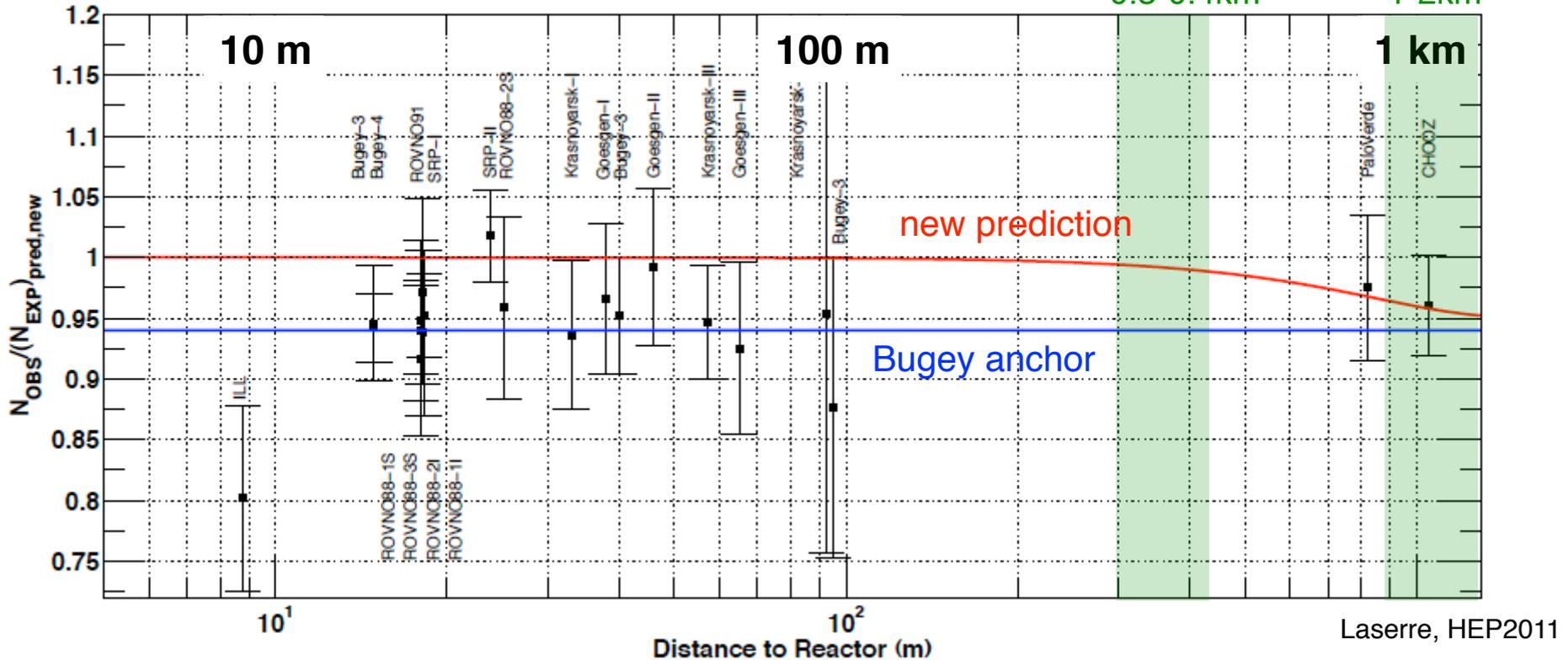


# Reactor Anomaly

## Reactor flux measurements vs distance

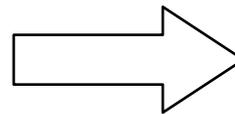
reactor  $\theta_{13}$   
near detector  
0.3-0.4km

reactor  $\theta_{13}$   
far detector  
1-2km



### updates in 2011

- new reactor antineutrino spectra
- re-analysis of 21 reactor results
- neutron lifetime correction, off-equilibrium effects



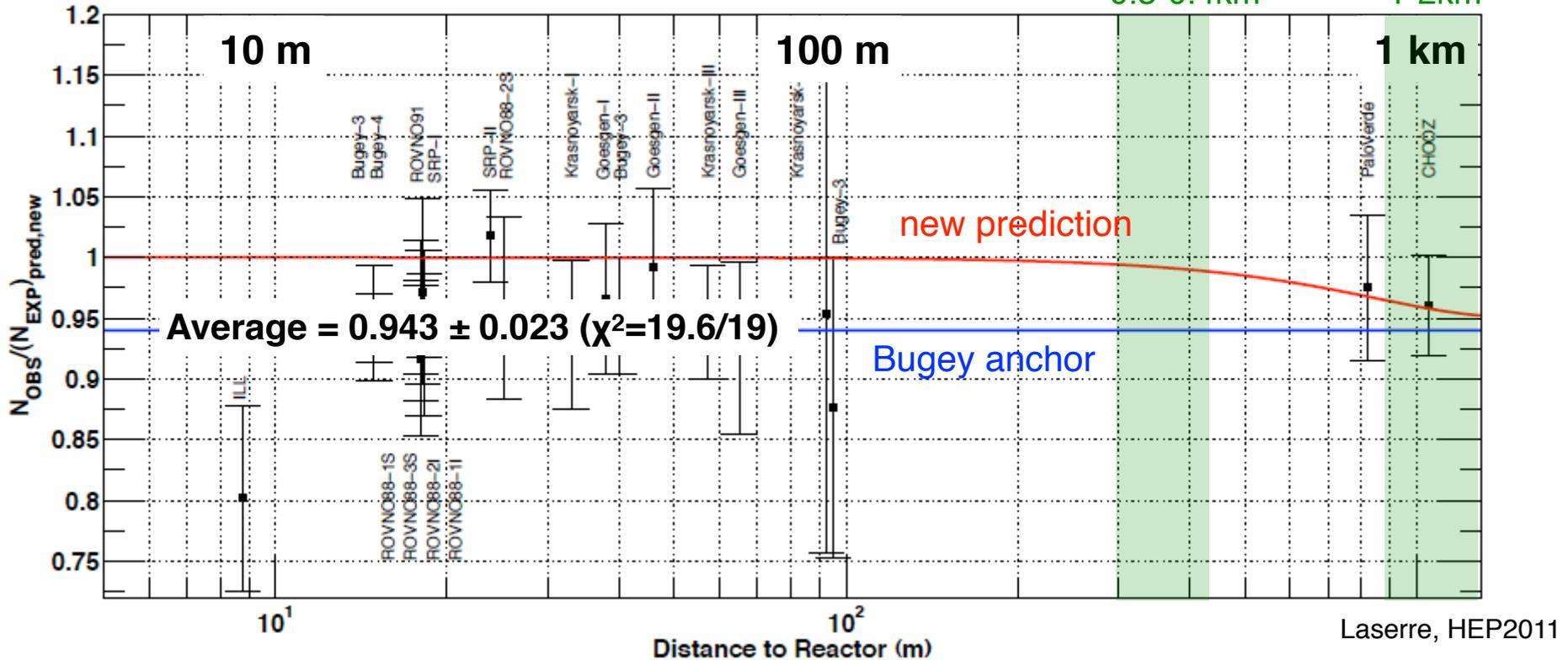
net 3% upward shift in energy averaged fluxes

# Reactor Anomaly

## Reactor flux measurements vs distance

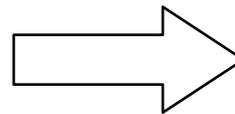
reactor  $\theta$   
near detector  
0.3-0.4km

reactor  $\theta$   
far detector  
1-2km



### updates in 2011

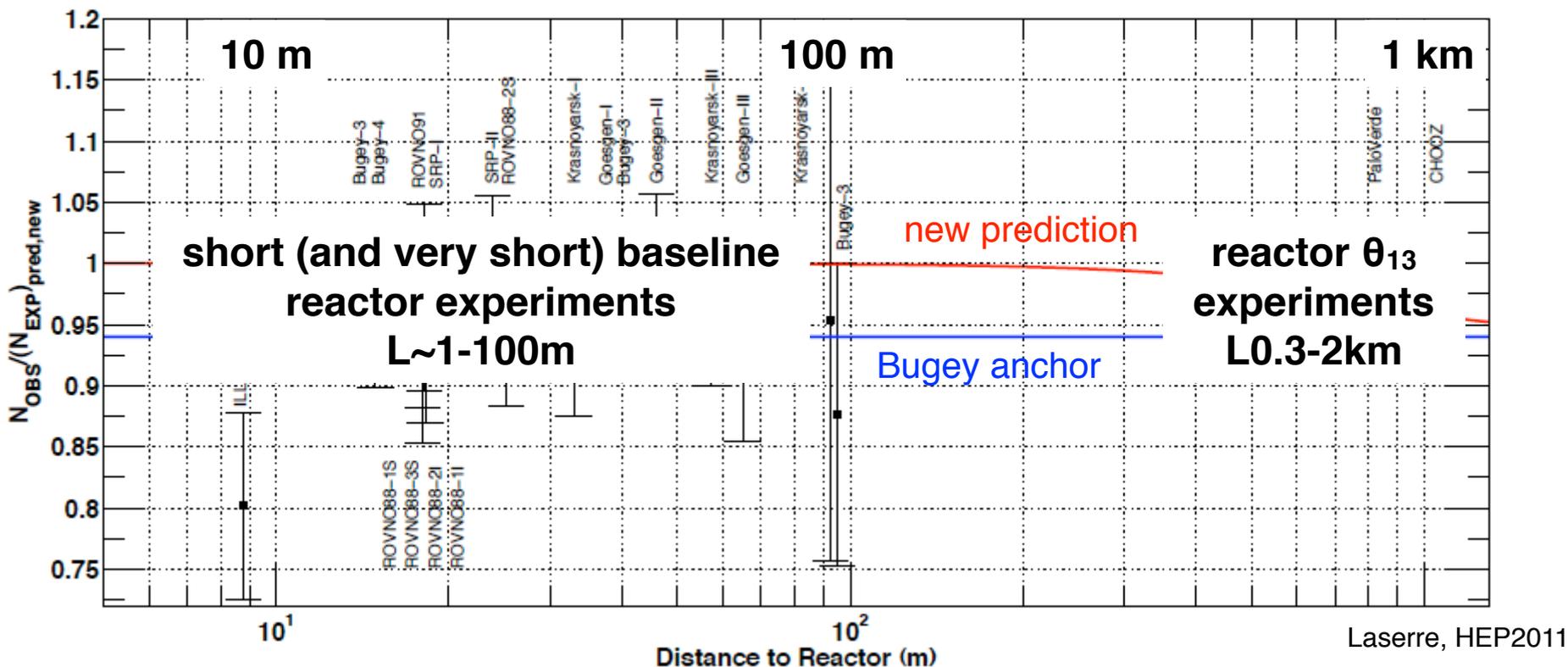
- new reactor antineutrino spectra
- re-analysis of 21 reactor results
- neutron lifetime correction, off-equilibrium effects



net 3% upward shift in energy averaged fluxes

# Reactor Anomaly

## Reactor flux measurements vs distance



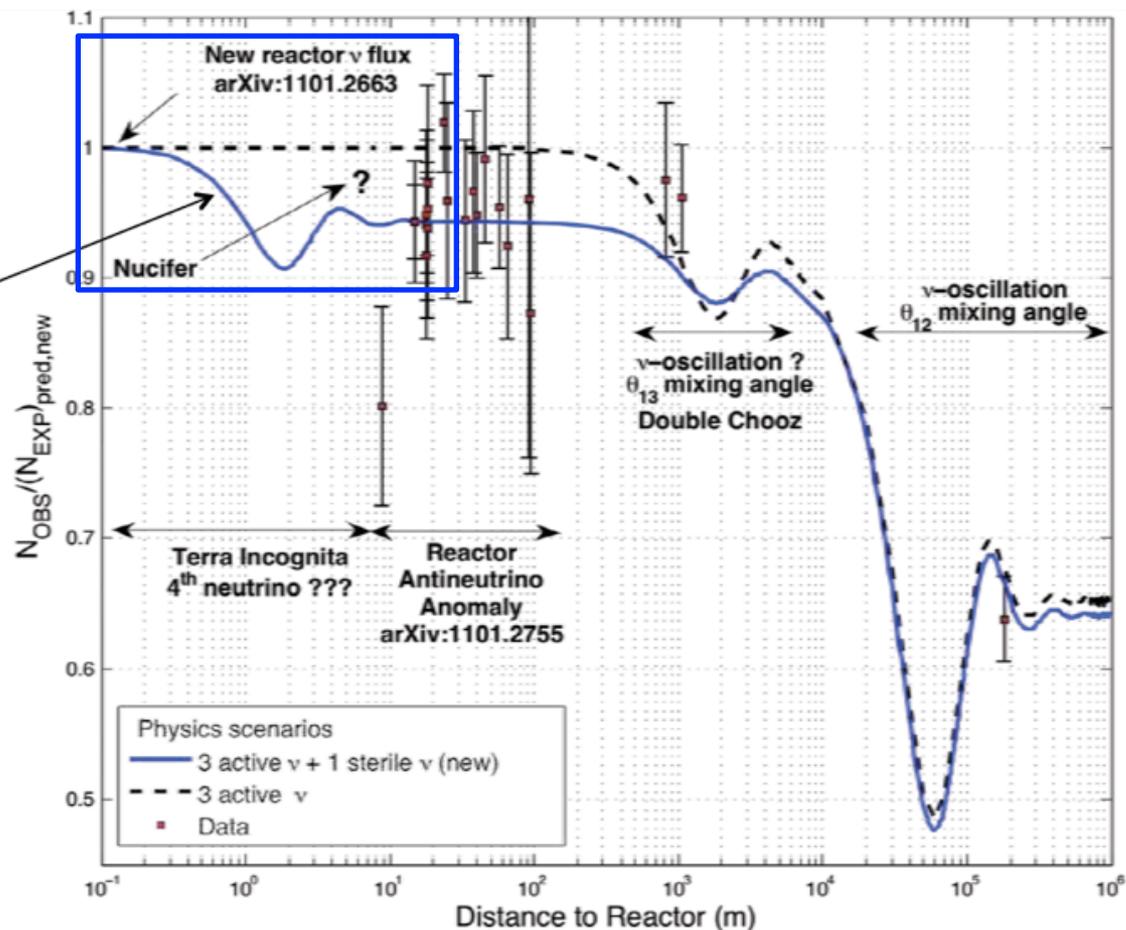
oscillation length of interest for sterile neutrino search: 2-3 m

reactor  $\theta_{13}$  experiments may help with understanding reactor flux normalization and spectrum but cannot provide proof/evidence of sterile  $\nu$

# Sterile $\nu$ Searches with Very Short Baselines: Reactors

Very short-baseline reactor experiments require  $L \sim 1-10\text{m}$

see Nathaniel's Bowden talk



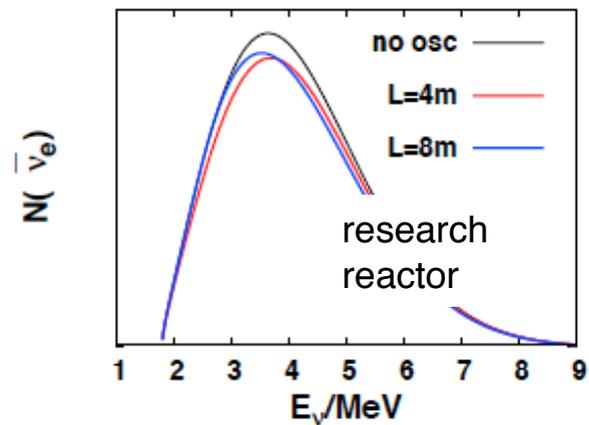
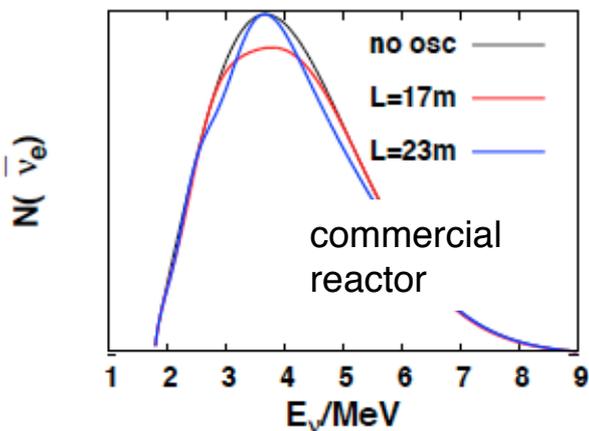
## Challenges:

- getting close to reactor core
- operating neutrino detector in high-background environment (typically very little overburden near reactor core, large gamma and neutron backgrounds)

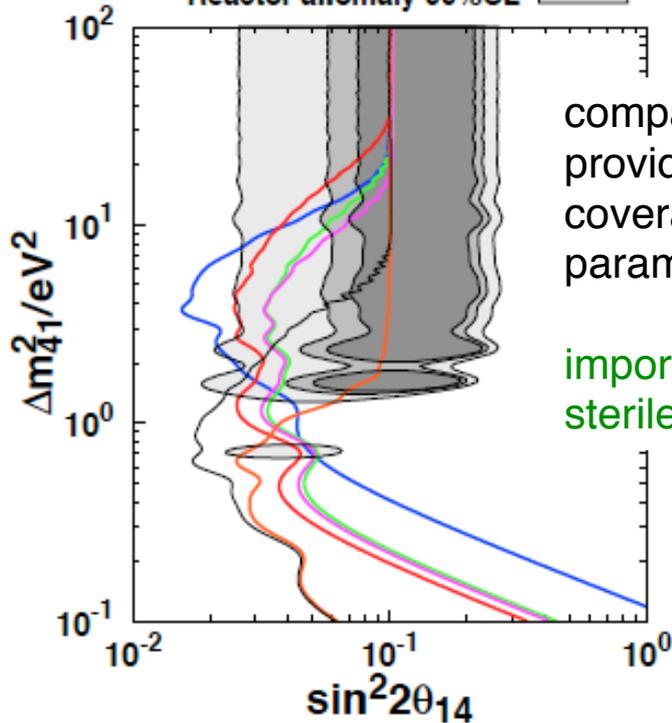
# Sterile $\nu$ Searches with Very Short Baselines: Reactors

Research vs commercial reactors and compact cores for very short-baseline reactor experiments at  $L \sim 1-10$  m

Energy spectrum w/ and w/out oscillation



- Joyo with  $L_F=4m, L_N=3m$  — blue line
- Joyo with  $L_F=8m, L_N=4m$  — red line
- ILL with  $L_F=8m, L_N=4m$  — green line
- Osiris with  $L_F=8m, L_N=4m$  — magenta line
- Commercial with  $L_F=23m, L_N=17m$  — orange line
- Point-like with  $L_F=23m, L_N=17m$  — grey line
- Reactor anomaly 90% CL — dark grey shaded region
- Reactor anomaly 95% CL — medium grey shaded region
- Reactor anomaly 99% CL — light grey shaded region



compact research reactors provide generally better coverage of anomaly parameter space

important for definitive sterile nu experiment

Yasuda, arXiv: 1110.2579

# Sterile $\nu$ Searches with Very Short Baselines: Sources

Alternative Approach: Place source near or inside detector and search for  $\nu_e$  or  $\bar{\nu}_e$  disappearance.

## Advantages

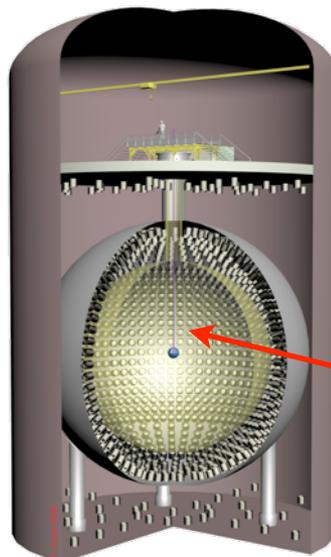
- baseline can be as short as needed
- detectors can be underground to minimize backgrounds
- potential for oscillometry (i.e. demonstrate oscillation signature vs baseline and energy)
- may be able to re-use existing, well-characterized detectors

## Challenges

- construct suitable, intense radioactive source
- regulatory and licensing requirements for radioactive source

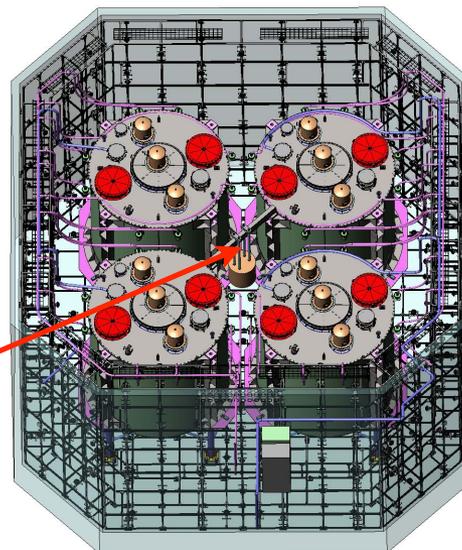
source inside detector

e.g. Ce-LAND



source next to detector

e.g. Daya Bay



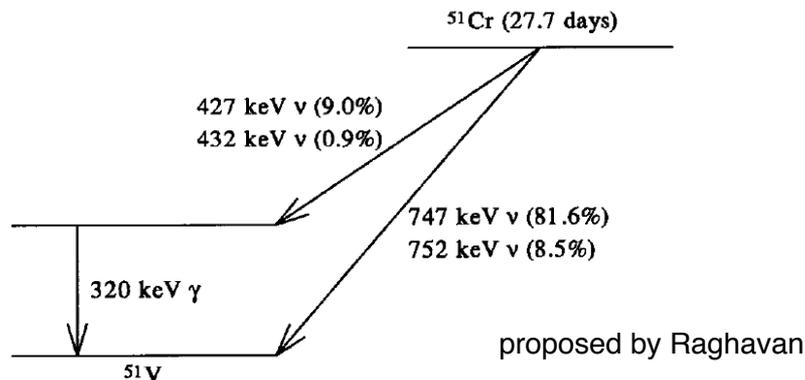
source

# Sterile $\nu$ Searches with Very Short Baselines: Sources

## A Variety of Sources and Detectors Are Feasible

Sources based on EC  
( $^{65}\text{Zn}$ ,  $^{51}\text{Cr}$ ,  $^{152}\text{Eu}$ ,  $^{37}\text{Ar}$ )

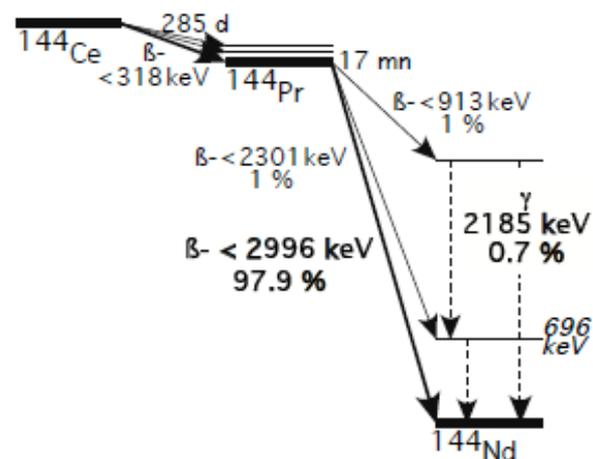
e.g.  $^{51}\text{Cr}$ , mono-energetic,  $\bar{\nu}_e$ , 750 keV



Decay scheme of  $^{51}\text{Cr}$  to  $^{51}\text{V}$  through electron capture.

Sources based on beta-decays

e.g.  $^{144}\text{Ce}$ - $^{144}\text{Pr}$ ,  $\bar{\nu}_e$ , continuous spectrum



arxiv:1107.2335  
Cribier et al

## Detection Channels & Proposed Experiments

Elastic Scattering: Borexino, SNO+Cr

Charged Current: LENS-Sterile, Baksan, Ce-LAND, Borexino, Daya Bay

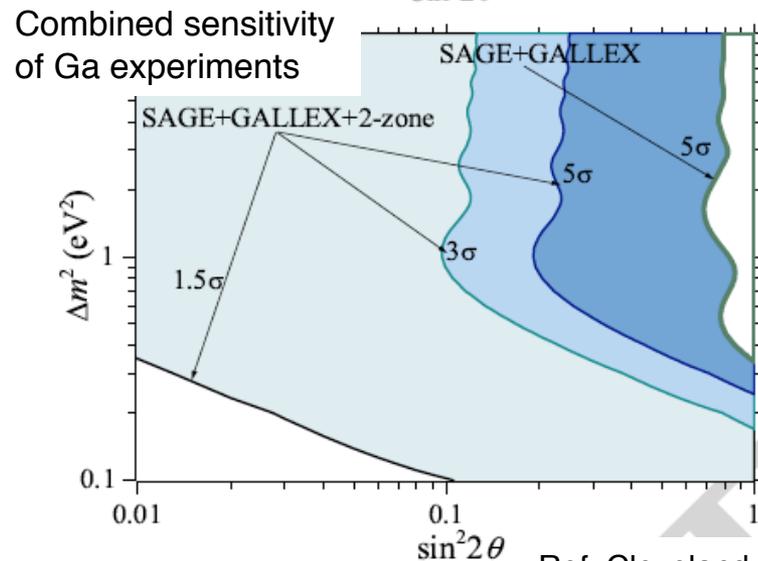
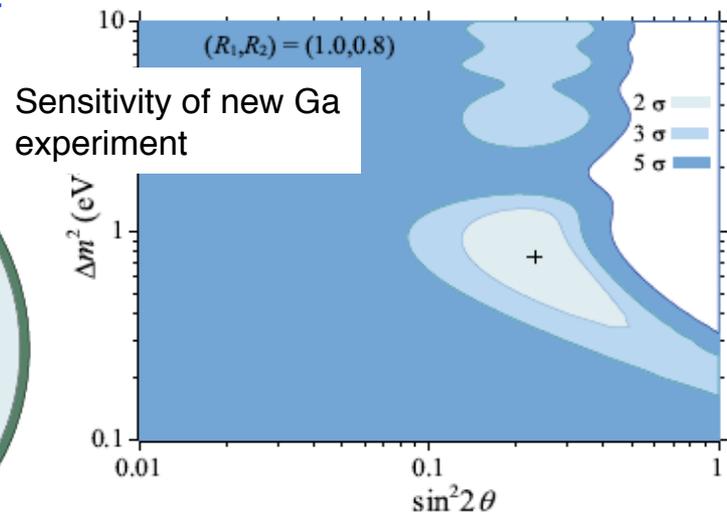
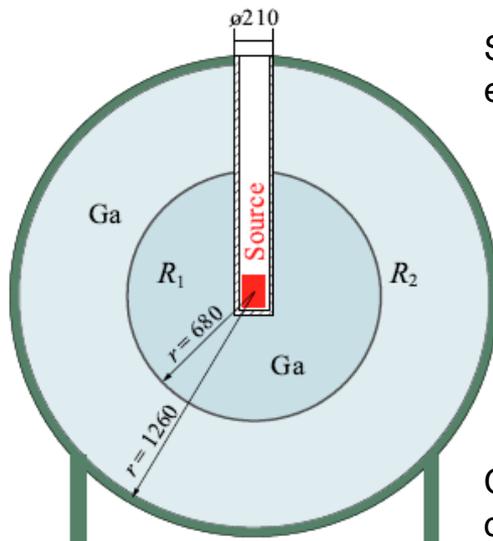
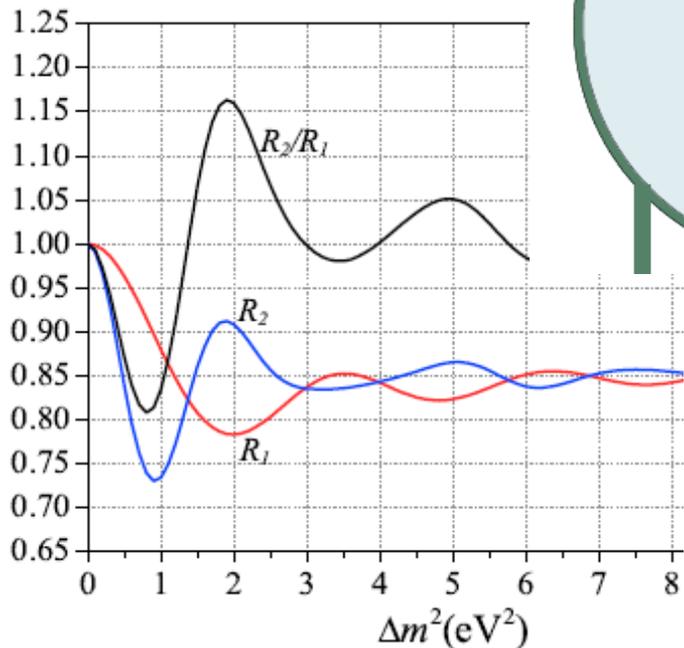
Neutral Current: RICOCHET

see following examples

# Short Baseline Search with Ga Target

## $^{51}\text{Cr}$ Source inside Dual Metallic Ga Target

measure ratio of capture rates in  $R_1$  and  $R_2$

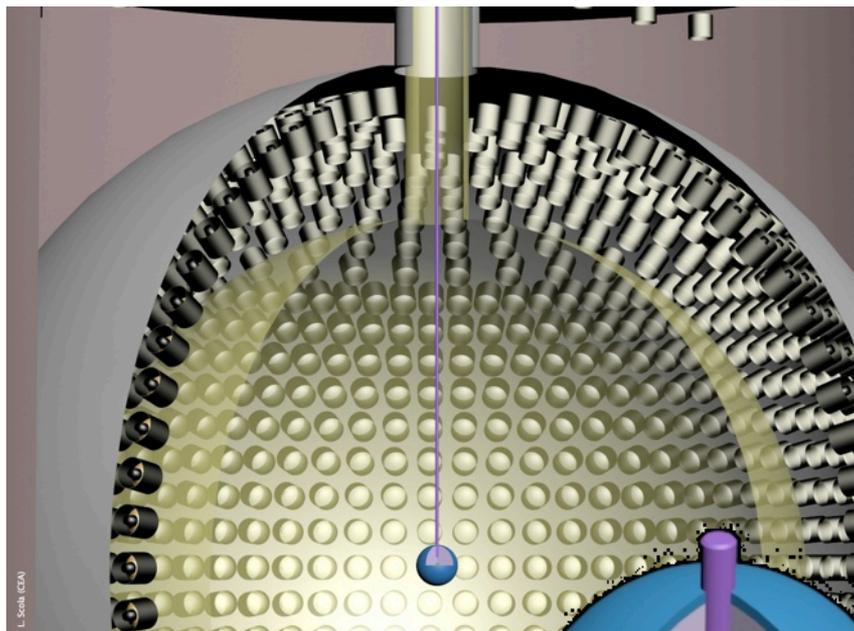


ratio of measured capture rates to predicted rate in inner and outer zones and their ratio  $R_2/R_1$

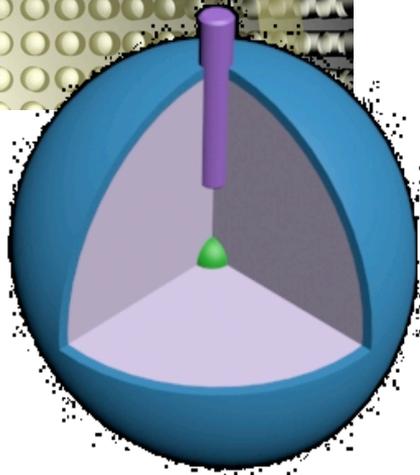
Ref: Cleveland et al.

# Ce-LAND

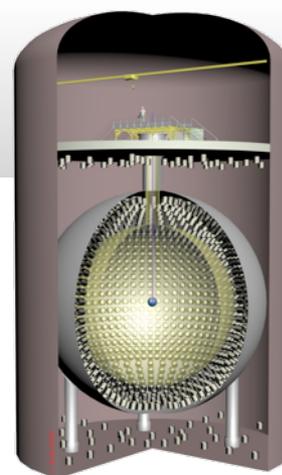
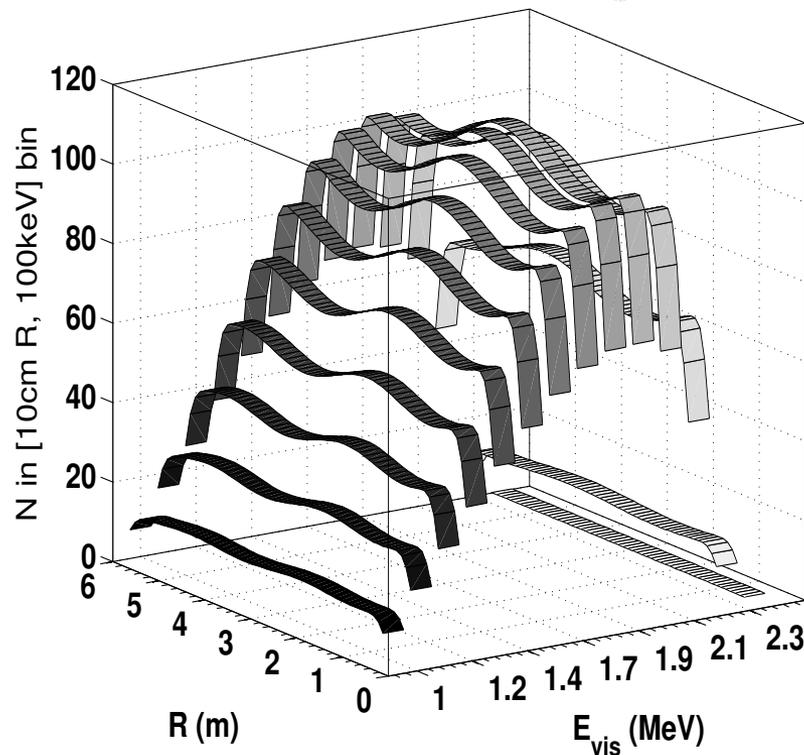
## $^{144}\text{Ce}$ source inside Liquid Scintillator Detector



Holder & cold finger  
R=8cm, 1.5 kg of Ce  
R=37 cm, Tungsten  
mass= 5 tons,  
diameter 82 cm  
teflon coating?



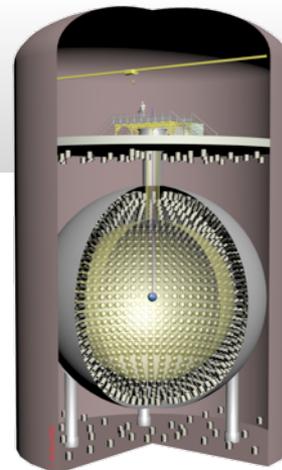
map oscillation effect in  
R and E



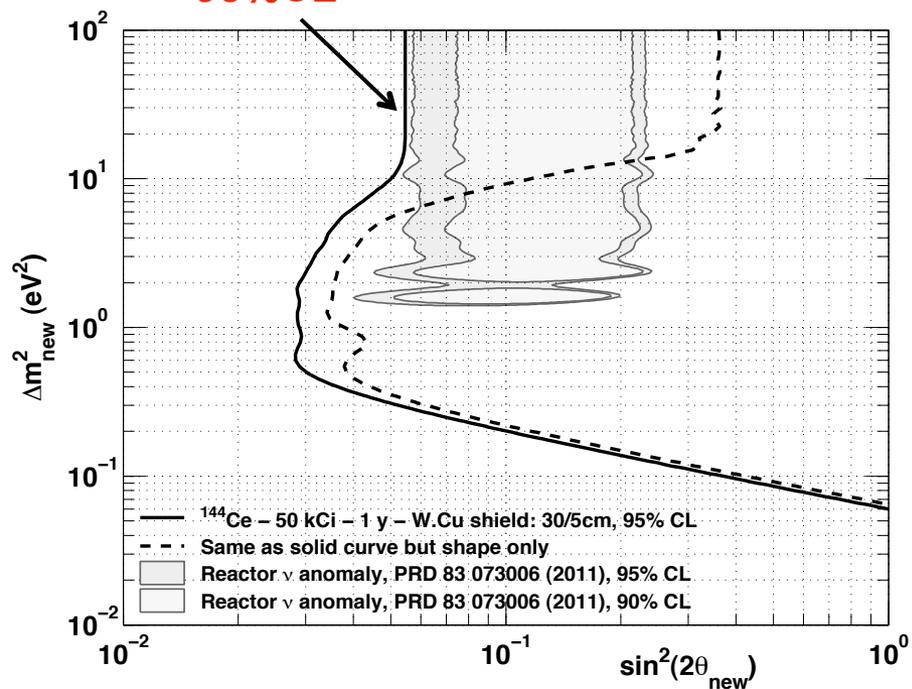
Ref: Lasserre

# Ce-LAND

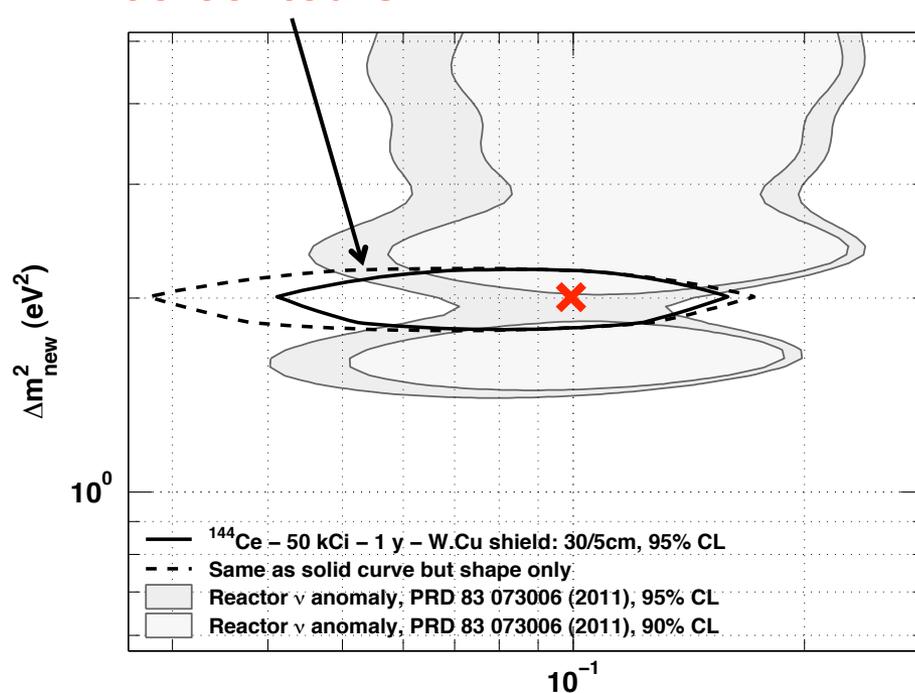
## Sensitivity and Discovery Potential



1 y – 50 kCi  
 $^{144}\text{Ce}-^{144}\text{Pr}$   
 95%CL



1 y – 50 kCi  
 $^{144}\text{Ce}-^{144}\text{Pr}$   
 5 $\sigma$  contours



$\sin^2 2\theta = 0.1$  &  $\Delta m^2 = 2 \text{ eV}^2$  tagged at 5 $\sigma$

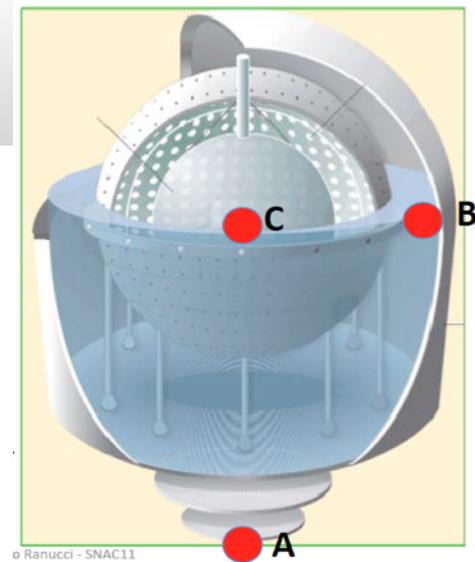
Ref: Lasserre

# Borexino Source Experiment

Source	decay	$\tau$ [days]	Energy [MeV]	Kg/MCi	W/kCi
$^{51}\text{Cr}$	e-capture ( $E_\gamma=0.32$ MeV 10%)	40	0.746 81%	0.011	0.19
$^{90}\text{Sr}-^{90}\text{Y}$	Fission product $\beta^-$	15160	<2.28 MeV 100%	7.25	6.7
$^{144}\text{Ce}-^{144}\text{Pr}$	Fission product $\beta^-$	411	<2.9975 MeV 97.9%	0.314	7.6

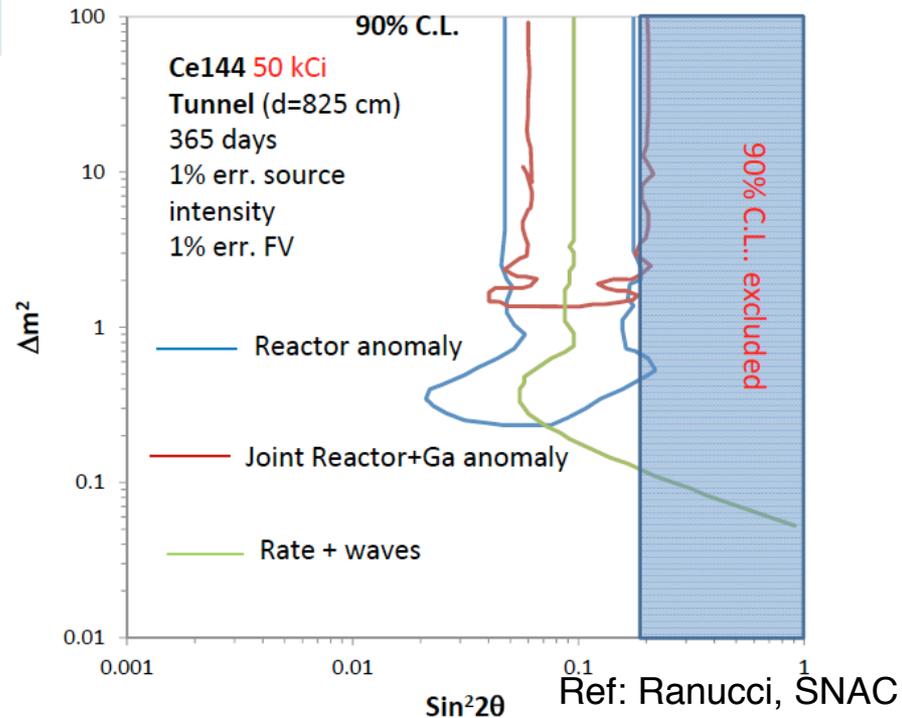
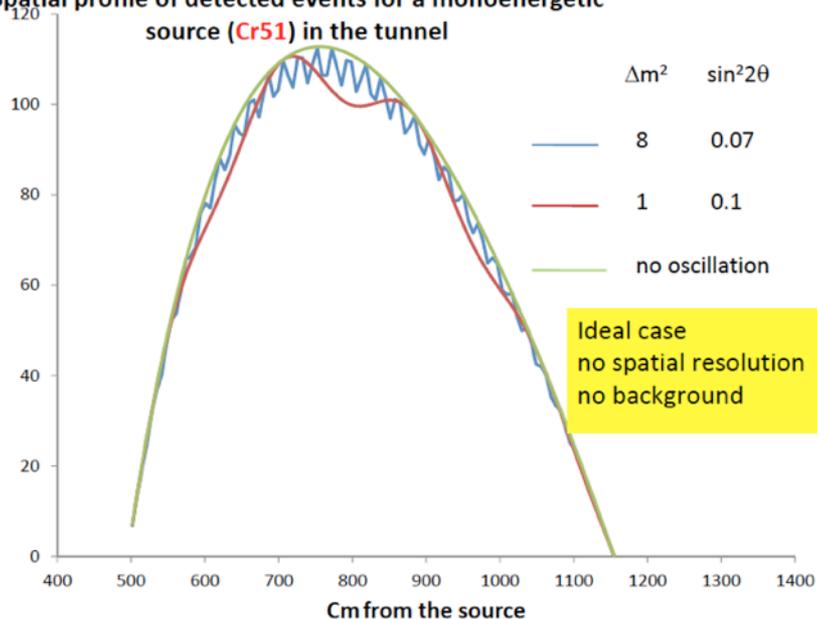
different sources and locations under consideration

inside and outside detector



© Ranucci - SNAC11

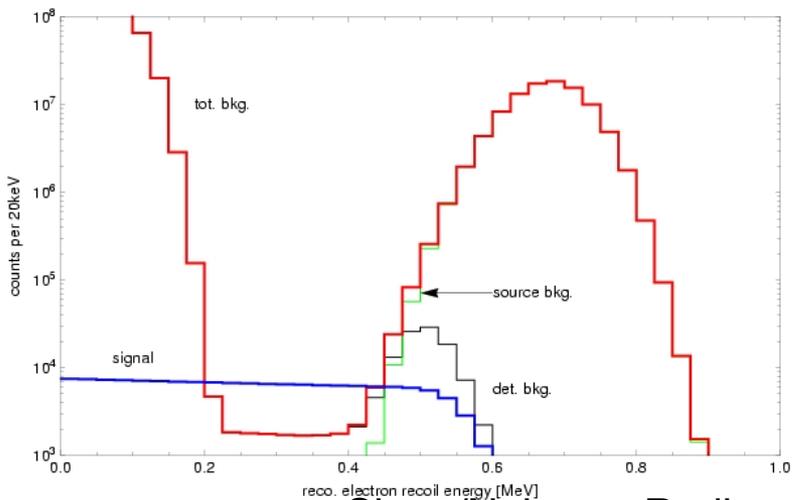
Spatial profile of detected events for a monoenergetic source ( $\text{Cr}51$ ) in the tunnel



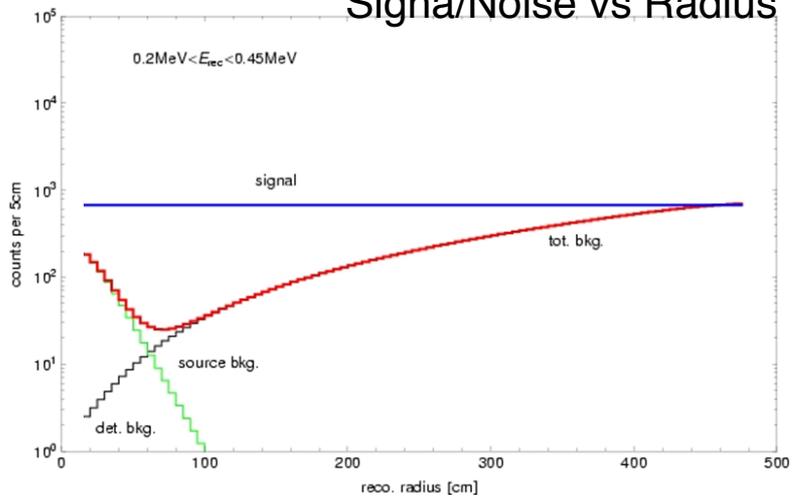
# SNO-Cr

## $^{51}\text{Cr}$ source inside SNO+

Central Source Location in SNO+

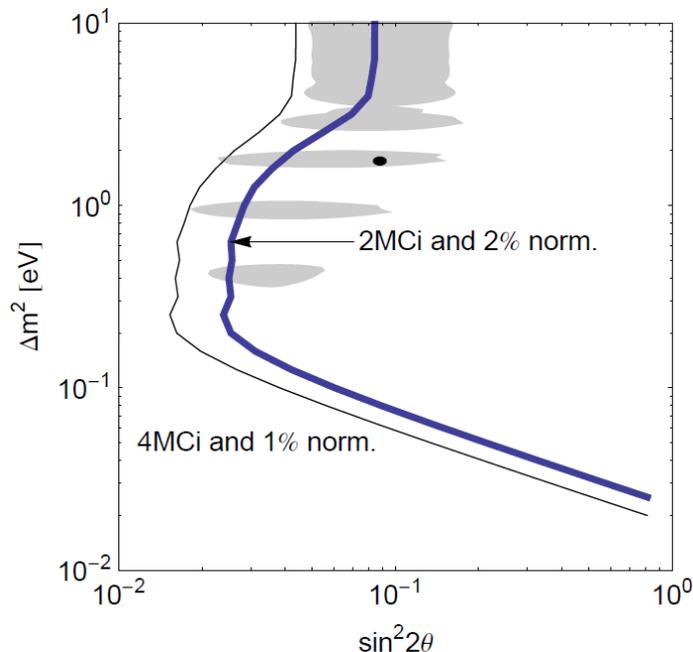
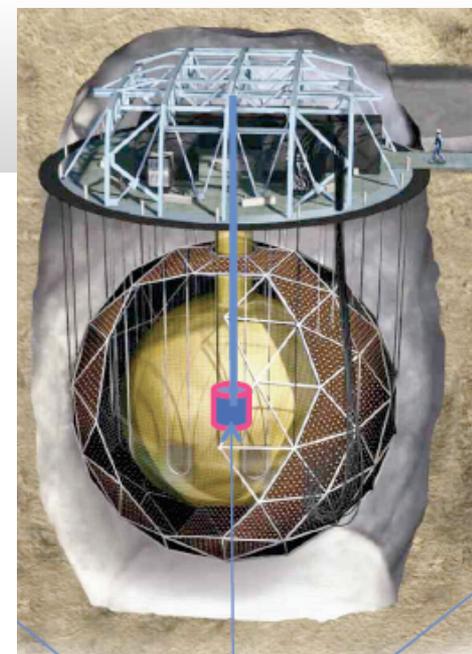


Signal/Noise vs Radius



SNO has widest neck/  
chimney of all liquid  
scintillator detectors

may be able to produce  
 $^{51}\text{Cr}$  source in US at High  
Flux isotope reactor at  
ORNL



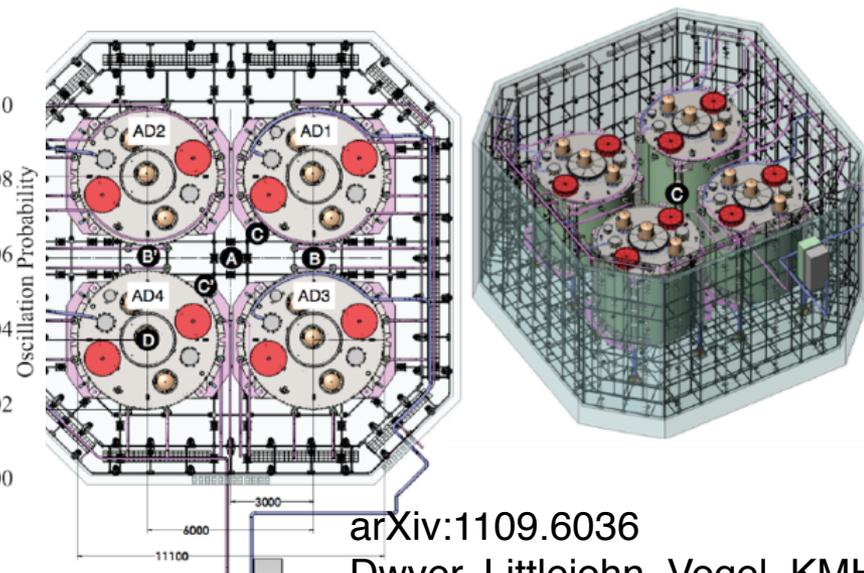
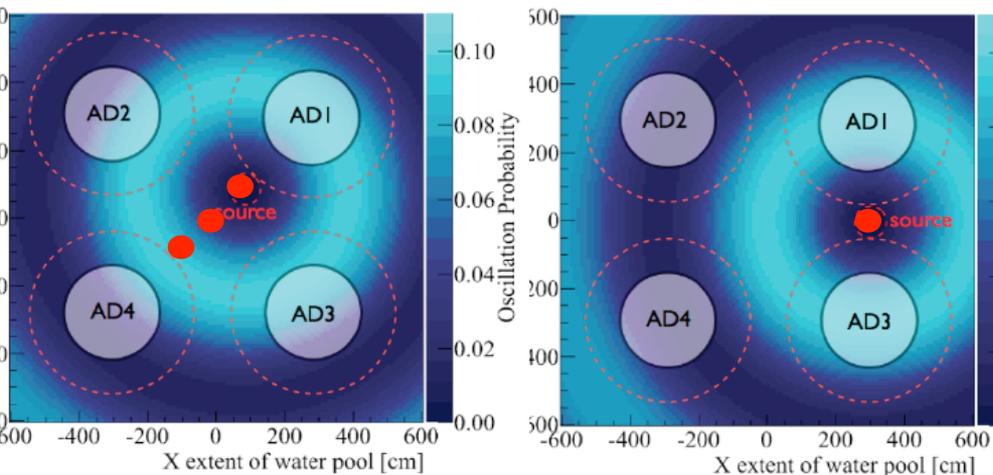
Ref: Link, Huber

# Daya Bay Sterile Neutrino Search

18 PBq  $^{144}\text{Ce}$  source at the Daya Bay far site

## Signal

- baseline range:  $\sim 1.5 - 8$  m
- $\nu$  energy range: 1.8 - 3 MeV
- 30k - 40k inverse beta-decay (IBD) events/per year
- Probing baselines from 1.5-8 m with an antineutrino source in the water pool of the Daya Bay Far Hall
- Advantageous to place source outside detectors in water pool.
- Multiple detectors allow for control of systematics.



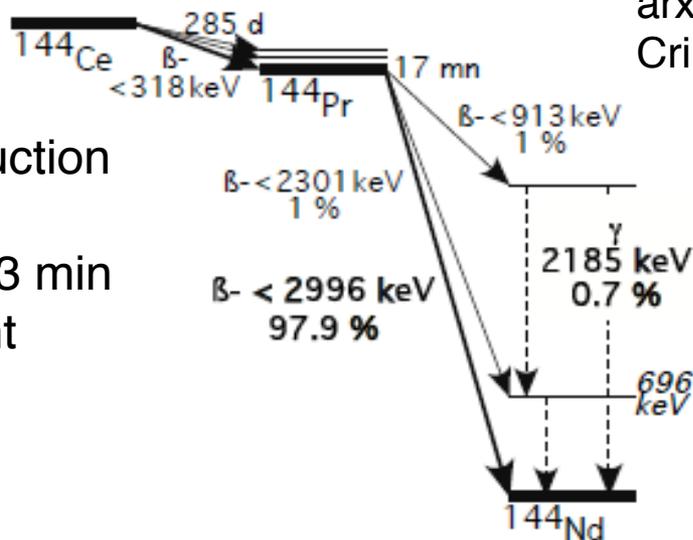
arXiv:1109.6036

Dwyer, Littlejohn, Vogel, KMH

# Daya Bay Sterile Neutrino Search

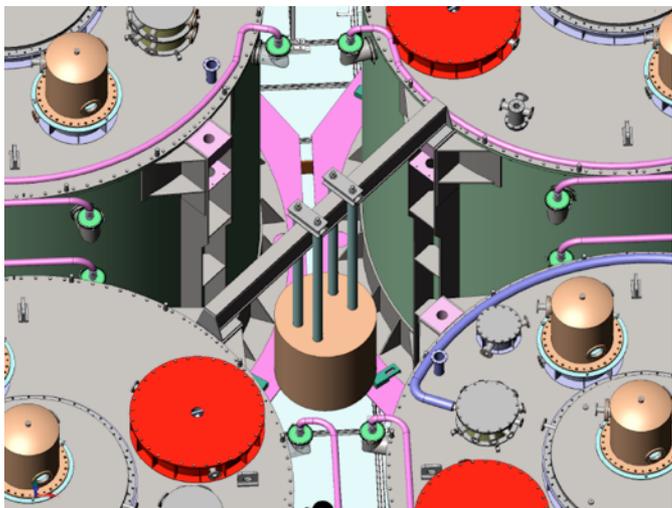
## $^{144}\text{Ce}$ - $^{144}\text{Pr}$ Antineutrino Source

- $Q_{\beta} > 1.8 \text{ MeV}$  (IBD threshold)
- lifetime long enough to allow for production and transport
- $T_{1/2} (^{144}\text{Ce}) = 285 \text{ days}$ ,  $T_{1/2} (^{144}\text{Pr}) = 17.3 \text{ min}$
- contained in fission fragments of spent nuclear fuel



arxiv:1107.2335  
Cribier et al

## Ce-144 source with W shielding



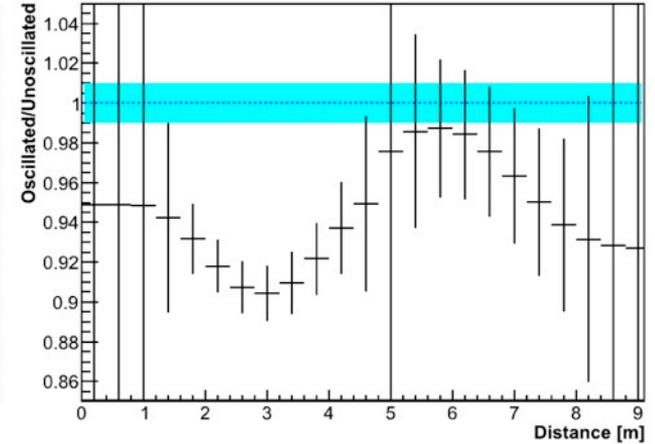
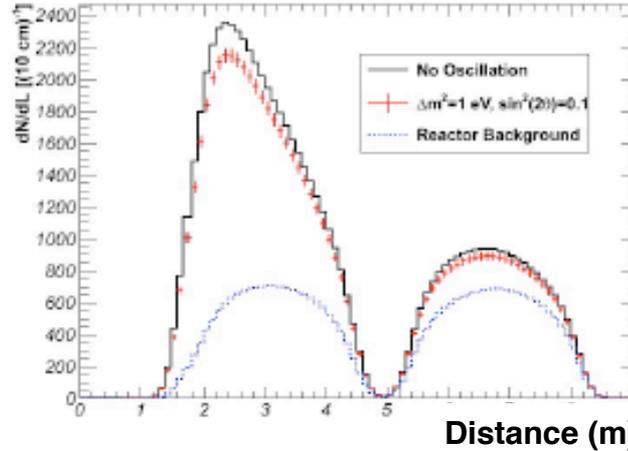
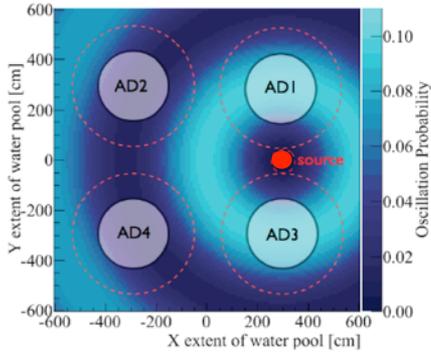
## Background

- $\sim 0.5 \text{ m}$  thick shielding makes source gammas negligible
- Water pool also shields, cools source outside detector
- Reactor neutrino 'background' well-known to  $< 1\%$  from near detectors

Littlejohn, KMH

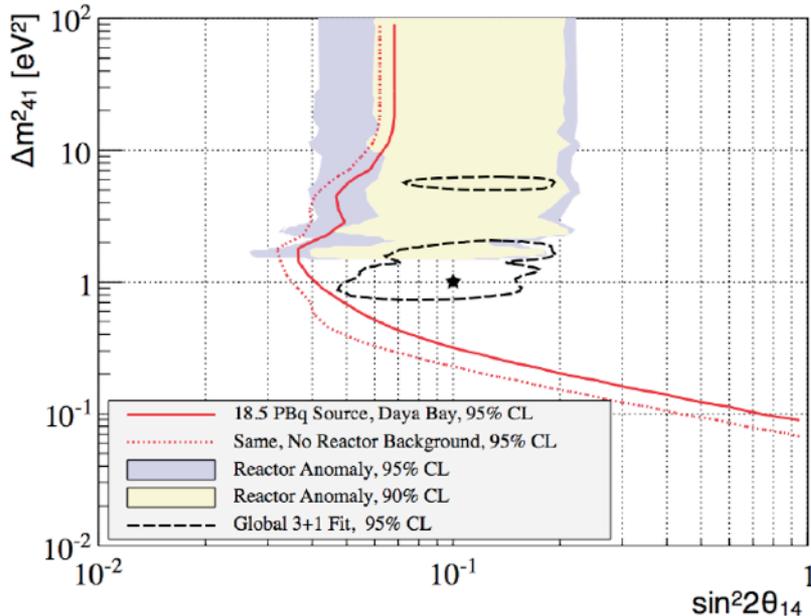
# Daya Bay Sterile Neutrino Search

## Signal and Event Distributions



arXiv:1109.6036

Dwyer, Littlejohn, Vogel, KMH



Sterile neutrino oscillations with mass  $> 1 \text{ eV}$  can be tested using  $^{144}\text{Ce}$  source in the Far Hall of the Daya Bay experiment after  $\theta_{13}$  measurement.

# Summary and Conclusions

- Reactor and source experiments provide a **complimentary way** to search for sterile neutrinos.
- Proposed experiments build on decades of experience with low-energy neutrino experiments, and **systematics are different from accelerator-based short-baseline experiments**.
- Energy and baseline-dependent signatures are critical for an unambiguous resolution of the current anomalies. Proposals need optimization to cover the entire region of anomaly parameter space.
- **Regulatory constraints for source production and logistics at reactors will likely constrain the number of feasible proposals.**
- There is no “fast” way to resolve the current anomalies. **Experiments need to be designed with sufficient systematic control, background mitigation, and physics reach to enable a  $>5\sigma$  discovery or exclusion**

More information in SNAC white paper: [http://cnp.phys.vt.edu/white\\_paper/](http://cnp.phys.vt.edu/white_paper/)

